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# Anomalous Dispersion Corrections Computed from Self-Consistent Field Relativistic Dirac-Slater Wave Functions\*

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Anomalous dispersion terms,  $\Delta f'$  and  $\Delta f''$ , have been computed for elements number 10 through 98 for five commonly used X-ray wavelengths. Oscillator strengths were calculated from self-consistent field relativistic Dirac wave functions which included in the potential Slater's approximate exchange contribution and Latter's self-interaction term.

### Introduction

In recent years, as counter devices have come into common use, it has been routinely possible to measure the effects of anomalous dispersion in X-ray diffraction experiments. The importance of this effect has been pointed out by Templeton (1955). Dauben & Templeton (1955) have published values of the anomalous dispersion corrections for most of the elements, as calculated by the method of Parratt & Hempstead (1954). An extension of Dauben & Templeton's (1955) calculations is given in International Tables for X-Ray Crystallography (1962). These calculations require a knowledge of the oscillator strengths and the manner in which the photoelectric absorption coefficient varies with the wavelength of the incident X-ray. An attempt has been made to improve the previous calculations by using atomic wave functions to obtain the oscillator strengths.

Initially, the Hartree wave functions recently computed by Boyd, Larson & Waber (1963) were used for calculation of the oscillator strengths. It was soon clear that, for heavy elements in particular, wave functions obtained by the simple Hartree model were completely inadequate for this purpose. The method was then checked with wolfram and uranium relativistic wave functions which had been calculated by Cohen (1960), and the results were considerably improved. Then, relativistic wave functions, which had been computed by Liberman, Waber & Cromer (1965), became available for all atoms. These relativistic wave functions include Slater's (1951, 1960) approximate exchange correction, and Latter's (1955) self-interaction term; hence we call them Dirac-Slater (DS) wave functions. The final set of oscillator strengths and dispersion terms in this paper was obtained from these DS wave functions.

### Theory

The theoretical treatment as given by James (1948) will be summarized. The total atomic scattering factor, f, is in general complex and is given by

$$f = f_0 + \Delta f' + i\Delta f'' \tag{1}$$

where  $f_0$  is the atomic scattering factor for radiation with frequency much higher than any absorption edge and  $\Delta f'$  and  $\Delta f''$  are the real and imaginary components of the anomalous dispersion. Neglecting damping, which is of importance only when the incident frequency is within about 10<sup>-4</sup> of an absorption edge (Parratt & Hempstead, 1954),  $\Delta f'$  is given by

$$\Delta f' = \sum_{k} \int_{\omega_k}^{\infty} \frac{\omega^2 (dg/d\omega)_k}{\omega_i^2 - \omega^2} d\omega$$
 (2)

where the summation is over all the absorption edges and

 $\omega_k =$  frequency of the absorption edge,

 $\omega_i = \text{incident frequency},$ 

- $(dg/d\omega)_k =$ oscillator density of type k at frequency  $\omega$ ,
- $(dg/d\omega)_k d\omega =$  the number of virtual oscillators of type k having frequencies between  $\omega$  and  $\omega + d\omega$ .

The imaginary component,  $\Delta f''$ , is given by

$$\Delta f^{\prime\prime} = \frac{1}{2} \pi \omega_i \sum_k (dg/d\omega)_{k_i} \tag{3}$$

where the summation is again over all of the absorption edges.

If we knew the functions  $(dg/d\omega)_k$  we could calculate  $\Delta f'$  and  $\Delta f''$  directly from equations (2) and (3). These functions are not known, although Hönl (1933) has made satisfactory calculations for the K electrons assuming them to be hydrogen-like. Hönl has also

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# ANOMALOUS DISPERSION CORRECTIONS

## Table 1. Calculation of oscillator strengths for uranium

		$g_{n'l'j'}^{n\ l\ j}  imes 1$	04	Coordina	te matrix ele	ment $\langle \mathbf{r} \rangle$	Wavelength of transition										
Transition	н	HD	DS	н	HD	DS	н	HD	DS	Observed							
$1s^{1} \rightarrow 2p^{1}$	1409	972	977	1.358 - 2*	$1 \cdot 116 - 2$	$1 \cdot 121 - 2$	0·1325 Å	0·1297 Å	0·1302 Å	0·1310 Å							
$2p^{\frac{3}{2}}$	2898	1924	1937	1.358 - 2	1.089 - 2	1.094 - 2	0.1288	0.1247	0.1251	0.1260							
$3p\frac{1}{2}$	225	127	126	5.027 - 3	3.736 - 3	3.727 - 3	0.1139	0.1113	0.1116	0.1123							
$3p^{\frac{3}{2}}$	452	338	340	5.027 - 3	$4 \cdot 294 - 3$	4.314 - 3	0.1133	0.1104	0.1107	0.1114							
4 2 4 2	53	29	29	2.398 - 3	1.761 - 3	1.759 - 3	0.1100	0.1075	0.1078	0.1084							
$4p\frac{3}{8}$	106	81	83	2.398 - 3	$2 \cdot 075 - 3$	$2 \cdot 101 - 3$	0.1099	0.1073	0.1076	0.1082							
$5p\frac{1}{5}$	12	7	7	1.153 - 3	8.408 - 4	8.469 - 4	0.1091	0.1066	0.1069 )	0.1074							
5p3	25	19	19	1.153 - 3	9.925 - 4	1.012 - 3	0.1091	0.1066	0.1068	0.1074							
$6n\frac{1}{5}$	2	1	1	$4 \cdot 432 - 4$	$3 \cdot 361 - 4$	3.541 - 4	0.1089	0.1065	0·1066								
$6p\frac{3}{2}$	3	3	3	$4 \cdot 432 - 4$	3.767 - 4	4.069 - 4	0.1089	0.1064	0.1066								
$2s\frac{1}{2} \rightarrow 2p\frac{1}{2}$	156	135	142	-6.050-2	-4.832-2	-4.806-2	23.7784	17.5177	16.4805								
$-2p\frac{3}{2}$	1924	2252	2263	-6.050-2	-5.505-2	-5.477 - 2	$3 \cdot 8523$	2.7245	2.6842								
$3p\frac{\overline{1}}{2}$	1260	1055	1082	$3 \cdot 120 - 2$	$2 \cdot 795 - 2$	$2 \cdot 823 - 2$	0.7824	0.7494	0.7462	0.7479							
$3p\frac{5}{3}$	2611	866	899	$3 \cdot 120 - 2$	1.744 - 2	1.773 - 2	0.7552	0.7118	0.7081	0.7102							
$4p\frac{1}{2}$	227	187	190	1.188 - 2	1.060 - 2	1.063 - 2	0.6300	0.6074	0.6038	0.6042							
$4p\frac{3}{2}$	456	190	198	1.188 - 2	7.510 - 3	7.642 - 3	0.6270	0.6007	0.5970	0.5988							
$5p\frac{1}{3}$	50	41	42	5.446 - 3	4.830 - 3	$4 \cdot 866 - 3$	0.6018	0.5798	0.5756	0.5766							
$5p\frac{2}{3}$	100	43	45	5.446 - 3	3.497 - 3	3.573 - 3	0.6017	0.5783	0.5741	0.5750							
$6p\frac{1}{2}$	7	6	7	$2 \cdot 076 - 3$	1.912 - 3	2.012 - 3	0.5968	0.5740	0.5696 )	0 5706							
$6prac{3}{2}$	15	6	7	$2 \cdot 076 - 3$	1.321 - 3	1.428 - 3	0.5968	0.5738	0.5694 ∫	0.9700							
$2p\frac{1}{2} \rightarrow 1s\frac{1}{2}$	- 1409	-972	- 977	1.358 - 2	1.116 - 2	$1 \cdot 121 - 2$	0.1325	0.1297	0.1302								
$2s\frac{1}{2}$	-156	-135	137	-6.050-2	-4.832 - 2	-4.723 - 2	23.7784	17.5177	16.4801								
$3s_{2}^{1}$	161	136	142	$1 \cdot 145 - 2$	1.036 - 2	1.060 - 2	0.8238	0.8006	0.7990	0.8052							
$3d\frac{3}{2}$	6945	3865	3893	5.074 - 2	3.704 - 2	3.710 - 2	0.7508	0.7188	0.7162	0.7200							
$4s\frac{1}{2}$	27	<b>23</b>	<b>24</b>	$4 \cdot 175 - 3$	3.825 - 3	3.888 - 3	0.6512	0.6343	0.6318	0.6355							
$4d\frac{3}{2}$	946	636	641	1.722 - 2	1.389 - 2	1.392 - 2	0.6351	0.6143	0.6117	0.6148							
$5s_2^1$	6	5	6	1.953 - 3	1.784 - 3	1.816 - 3	0.6189	0.6014	0.5982	0.6011							
$5d\frac{3}{2}$	165	112	116	7.079 - 3	5.741 - 3	5.830 - 3	0.6146	0.5955	0.5922	0.5947							
$6s\frac{1}{2}$	1	1	1	$8 \cdot 220 - 4$	7.714 - 4	8.078 - 4	0.6125	0.5940	0.5906								
$6d\frac{3}{2}$	2	1	2	1.488 - 3	1.048 - 3	1.581 - 3	0.6116	0.5927	0.5892	0.5920							
7 <i>8</i> 츛	0	0	0	$2 \cdot 230 - 4$	$2 \cdot 209 - 4$	2.622 - 4	0.6116	0.5927	0.5892								
$2p_2^3 \rightarrow 1s_2^1$	-1449	- 962	- 969	1.358 - 2	1.089 - 2	1.094 - 2	0.1288	0.1247	0.1251								
28 <u>5</u> 2.1	- 902	- 1120	- 1132	-0.030 - 2	-5.005-2	-0.195 - 9	3.0023 1.0026	1.0648	1.0641	1.0671							
	501	410 515	519	1.143 - 2	2.030 - 2	2.123 - 2	0.8079	0.0940	0.0999	0.0994							
345 245	5079	4704	4797	5.074 - 2	4.855 9	4.862 2	0.8001	0.0198	0.0000	0.0106							
$3u_{\overline{2}}$	0212	4704	4101	5.074-2	4.835 - 2	4.803-2	0.3301	0.9120	0.2022	0.0100							
$4s\frac{1}{2}$	23	59	59	$4 \cdot 175 - 3$	6.769 - 3	6.784 - 3	0.7591	0.7895	0.7868	0.7704							
4 <i>d</i> 2	82	70	69	1.722 - 2	1.015 - 2	1.006 - 2	0.7369	0.7588	0.4598	0.7540							
$4d_{2}^{2}$	734	681	682	1.722 - 2	1.082-2	1.080 - 2	0.7363	0.7569	0.7039	0.7940							
58±	5	13	13	1.953 - 3	$3 \cdot 077 - 3$	3.084 - 3	0.7152	0.7392	0.7004	0.1301							
$5d\frac{3}{2}$	14	12	12	7.079-3	0.033-3	0.008 - 3	0.7095	0.7900	0.70204	0.7263							
$5d\frac{2}{2}$	129	116	120	7.079 - 3	0.824 - 3	0.903 - 3	0.7094	0.7300	0.7200	0.7959							
58 <del>2</del>	1	2	3	$8 \cdot 220 - 4$	1.323 - 3	1.304-3	0.7007	0.7961	0.7910	0.7994							
$6a\frac{3}{2}$	0	U	U	1.488 - 3	1.189 - 3	1.1.4-3	0.7004	0.7961	0.7910	0.1224							
78 <u>‡</u>	0	0	0	z·230 - 4	3.181 - 4	4.421 - 4	0.7054	0.1201	0.1219								

\* The notation 1.358 - 2 means  $1.358 \times 10^{-2}$ .

applied his method to the L electrons but the approximation of hydrogen-like L electrons is not as good. Eisenlohr & Müller (1954) have also applied Hönl's method to the L electrons for several atoms.

It is possible to relate the functions  $(dg/d\omega)_k$  to the photoelectric absorption coefficients and to the oscillator strengths and, with certain assumptions regarding the absorption coefficients, perform the integration of equation (2). The total oscillator strength,  $g_K$ , for the K continuum is given by

$$g_K = \int_{\omega_K}^{\infty} (dg/d\omega)_K d\omega . \qquad (4)$$

The atomic photoelectric absorption coefficient  $\mu(\omega)$  is related to  $dg/d\omega$  by

$$(dg/d\omega) = (mc/2\pi^2 e^2)\mu(\omega)$$
(5)

where the symbols m, c and e have their usual meaning. Thus the oscillator density can be determined empirically from experimental values of  $\mu(\omega)$ .

The variation of  $\mu(\omega)$  for a particular edge is fairly well represented by the empirical formula

$$\mu(\omega) = \begin{cases} (\omega_k/\omega)^n \mu(\omega_k) & \text{for } \omega > \omega_k \\ 0 & \omega < \omega_k \end{cases}$$
(6)

where  $\omega_k$  is the frequency of the edge,  $\mu(\omega_k)$  is the absorption coefficient at the edge, and *n* has a value of the order of 3. Unfortunately, *n* varies, depending on the particular edge involved, and is a function of atomic number. Uncertainty in the value of *n* is a serious defect in the present analysis; further, the assumption made in equation (6) is not strictly valid.

If equations (5) and (6) are substituted into (4) we get

$$g_{k} = \int_{\omega_{k}}^{\infty} (mc/2\pi^{2}e^{2})\mu(\omega)d\omega$$
$$g_{k} = \frac{mc}{2\pi^{2}e^{2}}\mu(\omega_{k})\int_{\omega_{k}}^{\infty} \left(\frac{\omega_{k}}{\omega}\right)^{n}d\omega$$
$$g_{k} = \frac{mc}{2\pi^{2}e^{2}}\cdot\frac{\omega_{k}}{n-1}\mu(\omega_{k}).$$
(7)

If *n* and  $\mu(\omega_k)$  are determined from experiment,  $g_k$  can be calculated from (7). Another method of calculating  $g_k$  will be discussed below.

Equations (5) and (6) can be substituted into (2) resulting in

$$\Delta f'_{k} = \frac{mc}{2\pi^{2}e^{2}}\,\mu(\omega_{k})\omega_{k}^{n}\int_{\omega_{k}}^{\infty}\frac{\omega^{2}d\omega}{(\omega_{i}^{2}-\omega^{2})\omega^{n}}.$$
(8)

Equation (8) has been integrated for the general case by Parratt & Hempstead (1954). These authors give expressions for  $\Delta f'_k$  for various values of n as functions of  $g_k$  and x, where  $x = \omega_i / \omega_k$ .

Substitution of equations (5), (6) and (7) in (3) gives

$$\Delta f_k'' = \frac{mc}{4\pi e^2} \omega_i \left(\frac{\omega_k}{\omega_i}\right)^n \mu\left(\omega_k\right)$$
$$\Delta f_k'' = \frac{1}{2}\pi \frac{n-1}{x^{n-1}} g_k \tag{9}$$

where  $x = \omega_i / \omega_k$  as before. The dispersion terms for the whole atom are then obtained by using appropriate values of *n* and summing equations (8) and (9) over all the absorption edges.

Experimental values of  $\mu(\omega)$  are not available for all atoms and it is difficult to determine  $\mu(\omega)$  with accuracy. Thus there are few experimental values for oscillator strengths. It was therefore decided to calculate these quantities from the Hartree wave functions which have recently been computed.

Table 2. Oscillator strengths for wolfram and uranium

		Wol	fram		Uranium										
Edge	н	HD	DS	Hönl	н	HD	DS	Hönl							
$1s\frac{1}{2}$	1.04	1.24	1.24	1.12	0.96	1.30	1.30	1.12							
$2s\frac{\overline{1}}{2}$	0.88	1.13	1.10	1.35	0.64	1.04	1.03	1.28							
$2p\overline{\frac{1}{2}}$	0.89	1.25	1.24	0 ==	0.66	1.27	1.26	0.00							
$2p_2^{\overline{3}}$	2.31	2.33	2.31	3.00	2.18	$2 \cdot 20$	2.18	3.22							
$3s\frac{1}{2}$	0.83	1.01	1.01		0.51	0.87	0.87								
$3p\overline{\frac{1}{2}}$	1.15	1.36	1.33		0.78	1.22	1.19								
$3p^{\frac{2}{3}}$	3.15	3.02	2.98		2.93	2.82	2.74								
$3d\frac{3}{2}$	4.67	4.37	4.25		3.97	3.59	3.50								
$3d^{ ilde{5}}_{ ilde{2}}$	6.57	6.54	6.35		5.42	5.35	5.25								
4s	0.86	0.78	0.81		0.60	0.67	0.68								
$4p\frac{1}{3}$	0.89	0.88	0.91		0.69	0.80	0.82								
$4p\frac{3}{3}$	2.43	2.39	2.44		2.61	2.49	2.49								
$4d\frac{3}{8}$	3.17	2.80	2.88		4.18	3.42	2.43								
$4d^{\frac{5}{5}}$	4.45	<b>4</b> ·20	4.34		5.78	5.29	5.37								
$4f_{5}^{2}$	10.76	10.62	10.60		10.74	10.46	10.46								
$4f\frac{7}{2}$	14.56	14.47	14.48		14.13	1 <b>3</b> ·98	13.98								

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ANOMALOUS DISPERSION CORRECTIONS

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	$4f_{\overline{9}}^{7}$	t																							8.0033	13.848	15.476	17.162	19-232	23-455	26.944	30.382	31.732	33.507	35.714	37-541	39.388
	$4f_{\overline{2}}^{5}$	I																							8.3300	14-311	15-978	17-705	19.819	24.137	27-731	31.285	32.697	34.536	36.812	38.710	40-631
	$4d\frac{5}{2}$	T																						9-5771		34.450	36.655	38.920	41.573	46.973	51.658	56.312	58.279	60.673	$63 \cdot 503$	65.960	68-443
∕J <i>f</i> , ∕	$4d_2^3$																							9.9121		36-384	38.721	41.124	43-923	49.636	54.666	59.696	61.865	$64 \cdot 470$	61.519	70.205	72-926
'df' and	$4p_{\overline{2}}^{3}$																							17-616		51.775	54.495	57-250	60.475	926-99	72-807	78.648	81.228	84.246	87.711	90.815	93·958
lation of	$4p_2^1$																							19-475		62.054	$65 \cdot 483$	69.023	73.006	81.238	88.998	96.994	100.74	104.99	109.76	114.23	118-83
the calcu	$4s_2^1$																							24.114		71-465	75.148	78.948	83.194	91.966	100.28	108.86	112.90	117-45	122.53	127-32	132.23
, edges in orgs	$3d_{\overline{2}}^{\overline{5}}$																					19.617	49.785	75-998		197.27	204.72	212-31	220-39	239.90	253·04	269.48	277-47	285-97	294.99	303.72	312-55
<i>bsorption</i> are Rydb	$Edge 3d_2^3$																					119-911	50.778	77-849		206.00	213.96	222.07	230.69	248.36	265.76	283-56	292.27	301.52	311.32	320.85	330-53
<i>used as a</i> Units	$3p_2^3$									1.0603												31.518	68.098	98-966		241.14	249.75	258-55	267-85	286.90	305.69	324.90	334.32	344.28	354.81	365-07	375-48
snvalues :	$3prac{1}{2}$									I-0749												33.021	72.745	107-42		282.29	293-45	304.93	317.06	342.20	$367 \cdot 75$	<b>394·44</b>	407.89	422.10	437-09	452.05	467.40
lated eige	$3s\frac{1}{2}$									2.1259												39.503	82.488	119-46		303.60	315.26	327.23	339-86	366.04	392.65	420.45	434.48	449-27	464.85	480.42	496.39
5. Calcu	$2p_2^3$	1.4672	2.6622							18.160												197-08	350-93	474.17		1015.5	1044-7	1074-4	1105.0	1167-3	1230.5	1295.4	1328.0	1361-5	1395-8	1430-2	1465-0
Table	$2p_{\overline{2}}^{1}$	1.4758	2.6769							18-334												205.95	375.16	515.91		1197-9	1237-8	1278-7	1321.0	1408.6	1499-7	1595-3	1644.5	1695-3	1747-8	1801.3	1856-2
	$2s\frac{1}{2}$	3.1779	4.7397	6.5757	8.7498	11.137	13-754	16.607	19-701	23·038	27-282	31.909	36-331	40.972	45.846	94.871	03.06	11.66	20.65	30.04	39-83	22.45	98·19	43.53		1244·4	1285.3	1327-3	1370-6	1460-4	1553.8	1651.8	1702-3	1754-4	1808-2	1863-0	1919-3
	$1s_2^{I}$	63-077	78-203							233-55							I	Ĩ		Ι	-	1544·7 2	2540.1 3	3324·1 5		6872.2	1.067.1	7266.0	7469.1	7887.8	8322-8	8775-6	9.8006	9247-0	9490-7	9739-2	9993·1
	Atom	Ne	Na	$M_{g}$	AI 	ž	Ч	ŝ	C	Ar	K	Ca	$\mathbf{Sc}$	Ti	٧	Ga	Ge	As	Se	Br	$\mathbf{Kr}$	$\mathbf{T}_{\mathbf{c}}$	Xe	$\mathbf{Pm}^{:}$	Hg	$P_0$	$\mathbf{A}\mathbf{t}$	$\mathbf{Rn}$	$\mathbf{Fr}$	$\mathbf{Ac}$	$\mathbf{P}_{\mathbf{a}}$	Np	Pu	Am	Cm	Bk	IJ.

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The quantities g(k, n) are the quantum mechanical analogues of the classical oscillator strengths. These quantities are proportional to the transition probability of an electron going from state k to state nand thus are proportional to the coordinate matrix elements. The Thomas-Reiche-Kuhn sum rule states that for a one-electron atom,

$$\sum_{n} g(k, n) = 1$$
,

and it can be shown that for a many-electron atom the sum of the oscillator strengths is equal to Z, the atomic number.

Wheeler & Bearden (1934) applied this sum rule to obtain the oscillator strengths without the extrapolation of the continuum of the hydrogen-like atom used in the Hönl method. Recall, from equation (4), that  $g_k$  is given by the integral of the oscillator density from  $\omega_k$  to infinity, then this integral plus the transitions to all bound states is equal to unity by the sum rule. Therefore, for the K electrons,

$$g_K = 2\{1 - \sum_m g(K, m)\}$$
(10)

where g(K, m) is the oscillator strength of the virtual oscillator associated with the transition from the state K to the state m. The sum has to be taken over the relatively few occupied states. For other edges, the factor 2 in equation (10) is replaced by the multiplicity of the initial state. Wheeler & Bearden (1934) applied this method to the K edge only and to only a few atoms, there being few wave functions available at that time. This method has apparently not been again used until the present.

Bethe (1930) has given expressions for the oscillator strengths of hydrogen-like atoms for transitions  $nl \rightarrow n'l'$ . These expressions depend on the energy difference and transition probability. The Hartree wave functions could be used directly with Bethe's formulae but it is in principle more accurate to correct the Hartree energy levels by applying spin and relativistic corrections. These corrections split the p, d, and f states into two levels characterized by the three quantum numbers nlj, with the new levels having multiplicities (2j+1) where  $j=l\pm \frac{1}{2}$ .

Bethe's formulae for the oscillator strengths can no longer be used after the relativity and spin corrections have been applied. Dr James Young, of this Laboratory, has derived the expression for the oscillator strength of the transition from the state nlj to the state n'l'j'. Young's formula is given below.

$$g_{n\,l\,j}^{n'l'j} = \frac{1}{3} \varDelta E \max(l,\,l')(2j'+1) [W(l,\,j,\,l',\,j'\,;\,\frac{1}{2},\,1)]^2 \\ \times F' \left[ \int_0^\infty P_{n'l'}^* r P_{nl} dr \right]^2.$$
(11)

and final states, the factor (2j'+1) gives the multiplicity of the final state, W is the Racah coefficient (Racah, 1942), F' is the fractional occupancy of the final state and P is r times the Hartree radial wave function. The integral is the coordinate matrix element  $\langle n'l'|r|nl \rangle$  and its square is proportional to the probability of spontaneous transition from the state nl to the state n'l'.

When the relativistic wave functions described by Liberman, Waber & Cromer (1965) are used, the integral in equation (11) is replaced by

$$\int_{0}^{\infty} \left[ A_{nlj}^{*} A_{n'l'j'} + (-1)^{l-l'+j-j'} B_{nlj}^{*} B_{n'l'j'} \right] dr \quad (12)$$

where A and B are, respectively, r times the major and minor components of the wave functions.

### Initial calculations of oscillator strengths

The K oscillator strength for caesium was initially computed from the Hartree (H) wave functions, by using equations (10) and (11), in order to test the method and the computer program. The results essentially duplicated those of Wheeler & Bearden (1934). Spin and relativity corrections (Slater, 1960) were applied to the H eigenvalues. Differences between computed eigenvalues were used to obtain  $\Delta E$  for use in (11) rather than using experimental values of  $\Delta E$ . This was done for convenience and also because experimental values of  $\Delta E$  are not available for all transitions of interest for all elements.

More extensive oscillator strength calculations were then made for wolfram and uranium. The results for these heavy atoms did not agree well with Hönl's (1933) values and it was clear that the H wave functions were inadequate for this purpose. Relativistic Hartree-Dirac (HD) wave functions without exchange were available for wolfram and uranium (Cohen, 1960). The calculations were repeated with these wave functions and the results were considerably improved.

As an example, values of the individual terms used to calculate the oscillator strengths of the four innermost absorption edges of uranium are given in Table 1. The values from the H, the HD and the DS wave functions (DS wave functions were finally used for all atoms) are given so that the effect of better approximations to the true wave functions can be seen. The energy differences in Table 1 are given in terms of wavelength. The Hartree  $\Delta E$  values for many transitions differ appreciably from experimental values but it can be seen that this error is minor as compared with the error in the transition probabilities computed from these wave functions.

Table 2 lists, for wolfram and uranium, the oscillator strengths calculated from the three different wave functions. Hönl's (1933) values are also listed for comparison.

### Oscillator strengths and dispersion terms for elements 10 through 98

At this stage of the investigation the DS wave functions became available for all the atoms. Oscillator strengths were computed for elements 10 through 98 and these values are listed in Table 3. For elements 10 through 17, four absorption edges were considered; for elements 18 through 28, seven edges; elements 30 through 54, nine edges; elements 55 through 70, fourteen edges and for elements 71 through 98, sixteen edges were considered. Note that for neon and argon, for which the total number of electrons in the atoms is included in the edges considered, the sum of the oscillator strengths equals the number of electrons; thus the sum rule is satisfied.

The dispersion terms  $\Delta f'$  were then computed by using Parratt & Hempstead's (1954) solution of equation (8) and summing over the various edges considered. For all atoms, n=11/4 was used for the  $1s_{\frac{1}{2}}$  edge, n=7/3 for the  $2s_{\frac{1}{2}}$  edge and for all other edges, n=5/2 was used. These are the same values of *n* used by Dauben & Templeton (1955). The dispersion terms  $\Delta f''$  were computed from equation (9), again by summing over the various edges. Calculations were made for five different radiations in relatively common use. The results of these calculations are given in Table 4. Values are given only for  $\sin \theta/\lambda = 0$ . Although the dispersion terms are functions of  $\sin \theta/\lambda$ , the variation is so small that it can be neglected in most crystal structure work.

The values of  $\omega_k$ , the frequencies of the various absorption edges used in equations (8) and (9), were for the most part taken from the tabulation of energy levels by Sandström (1957). Where Sandström does not list energy levels, computed eigenvalues were used for  $\omega_k$ . The asterisk after an element in Table 4 indicates that one or more computed eigenvalues were used in the calculations. Table 5 gives the computed eigenvalues that were used.

These dispersion terms are, in general, not much different from those calculated by Dauben & Templeton (1955), because the oscillator strengths used here are not much different from those previously determined by other methods. However, the dispersion terms given here are probably as accurate as can be computed by the present theoretical methods, because reasonably accurate oscillator strengths have been used and because the contributions of the various absorption edges have been individually considered. More accurate wave functions, such as relativistic Hartree–Fock wave functions or those wave functions with correlation effects included, would probably not change the oscillator strengths very much.

The main limitations on the accuracy of these calculations are the uncertainty in the value of n in equations (8) and (9) and, more fundamentally, the questionable validity of equation (6), in which the exponential factor n is introduced. Also, for incident frequencies near an absorption edge, the exact position of the edge becomes of primary importance.

Sufficient information is given in Tables 3 and 5 so that one can compute the dispersion terms for other radiations or with other values of n for any atom.

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